## Modeling the Chemistry and Spectra of Water World Atmospheres with PICASO



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#### Abstract

So far, the quantitative study of **water world** atmospheres has lagged far behind the study of giant planets, sub-Neptunes, and terrestrial planets, which we aim to rectify. Here, we investigate water worlds' time evolving atmospheric chemistry, coupled to planetary cooling, mediated by the important and changing interface between a water layer and the magmatic rocky layer beneath. We use these chemistry calculations and corresponding fully coupled 1D radiative-convective atmospheric structure calculations to model spectral diagnostics of main atmospheric absorbers. A goal is to understand important absorbers beyond steam, which may be diagnostics of atmospheric and **atmosphere/mantle** interface conditions. This work enhances the capabilities of the open-source atmosphere modeling code **PICASO** to include atmospheres not dominated by H/He.

#### **The PICASO Climate Model**

 PICASO is a python-based open source atmospheric model with its legacy from the EGP model. It was developed and optimized to model the atmospheres of H<sub>2</sub>/He rich atmospheres, and has been used to model transiting gas giants, directly imaged planets, and brown dwarfs.

#### Expected Chemistry of Water Worlds

- To examine the abundances of molecules beyond H<sub>2</sub>O and CO<sub>2</sub>, we calculated the chemistry for a CI chondrite (bodies with nonvolatile chemical composition closest to that of the Sun) with 16% water by weight at a variety of temperatures and pressures
- We show the chemistry at two temperatures based on observations done by the COMPASS team with JWST

Why We Need to Create Models for Water World Atmospheres

- What **PICASO** can do:
- Create 1D radiative-convective thermochemical equilibrium models
- Model disequilibrium chemistry self-consistently
- The user can perform further analysis on the models by:
  - Adding clouds in post-processing
  - Using another code to perform a grid-retrieval on the models









# With the launch of JWST, we are now able to better study smaller planets, but they are still underrepresented in analysis



Graph of exoplanets as of 2021 that have spectroscopic measurements. Smaller, rocky planets make up the majority, but are understudied. Image adapted from Dmitry Savransky with data from NASA Exoplanet Archive

We can study the atmospheres of these planets through retrievals (allowing each atmosphere parameter to vary freely until one achieves a best fit model) or Schematic of the workflow of PICASO from Mukherjee et al 2023 ApJ 942 71

#### • How **PICASO** works:

- Inputs:
  - Stellar and planetary parameters
  - Initial guesses for the T(P) profile and the radiative
    - convective boundary
  - Atmospheric metallicity and C/O ratio
- Outputs:
  - A converged T(P) profile
  - Chemical abundances
  - Spectra transmission, thermal emission, and reflection

Chemical abundances for model atmospheres with a temperature of 550 K with 9 species present (above) and 1000 K with 13 species present (below)

 We find a wide range of molecules beyond H<sub>2</sub>O and CO<sub>2</sub> present in significant abundance. Clearly, more sophisticated models of water world atmospheres are needed

### **Next Steps**

- Making edits to PICASO in order to more accurately model water atmospheres
  - The user will set abundances for any combination of species including CO, H<sub>2</sub>O, CH<sub>4</sub>, NH<sub>3</sub>, CO<sub>2</sub>, N<sub>2</sub>, HCN, H<sub>2</sub>, PH<sub>3</sub>, C<sub>2</sub>H<sub>2</sub>, and Na and PICASO will

grid-retrievals (creating a grid of radiative-equilibrium models over a range of parameters and performing analysis on them). Grids can be modeled with **PICASO** using precomputed chemistries.



Scan to learn more about PICASO!



Contact us if you have questions! Anna: <u>akgagneb@ucsc.edu</u> Sagnick: <u>samukher@ucsc.edu</u> Jonathan: <u>ifortney@ucsc.edu</u> calculate the opacities on the fly

- $\circ~\mbox{Model}~\mbox{H}_2\mbox{O-H}_2\mbox{O}$  collision induced absorption in the atmosphere
- Use a  $H_2O/CO_2$  mixture as the background gas for calculating gaseous opacities instead of  $H_2/He$
- Extend the atmosphere down to the surface of the planet
  - Model interactions between a magma surface and a steam atmosphere as this may change the structure of the lower atmosphere
- Ultimate goal:
  - Perform analysis on a water world through all stages
    - Data reduction through COMPASS
    - Atmospheric models using PICASO