

## High-Resolution Infrared Transmission Spectroscopy

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### Observations and cross-correlation results

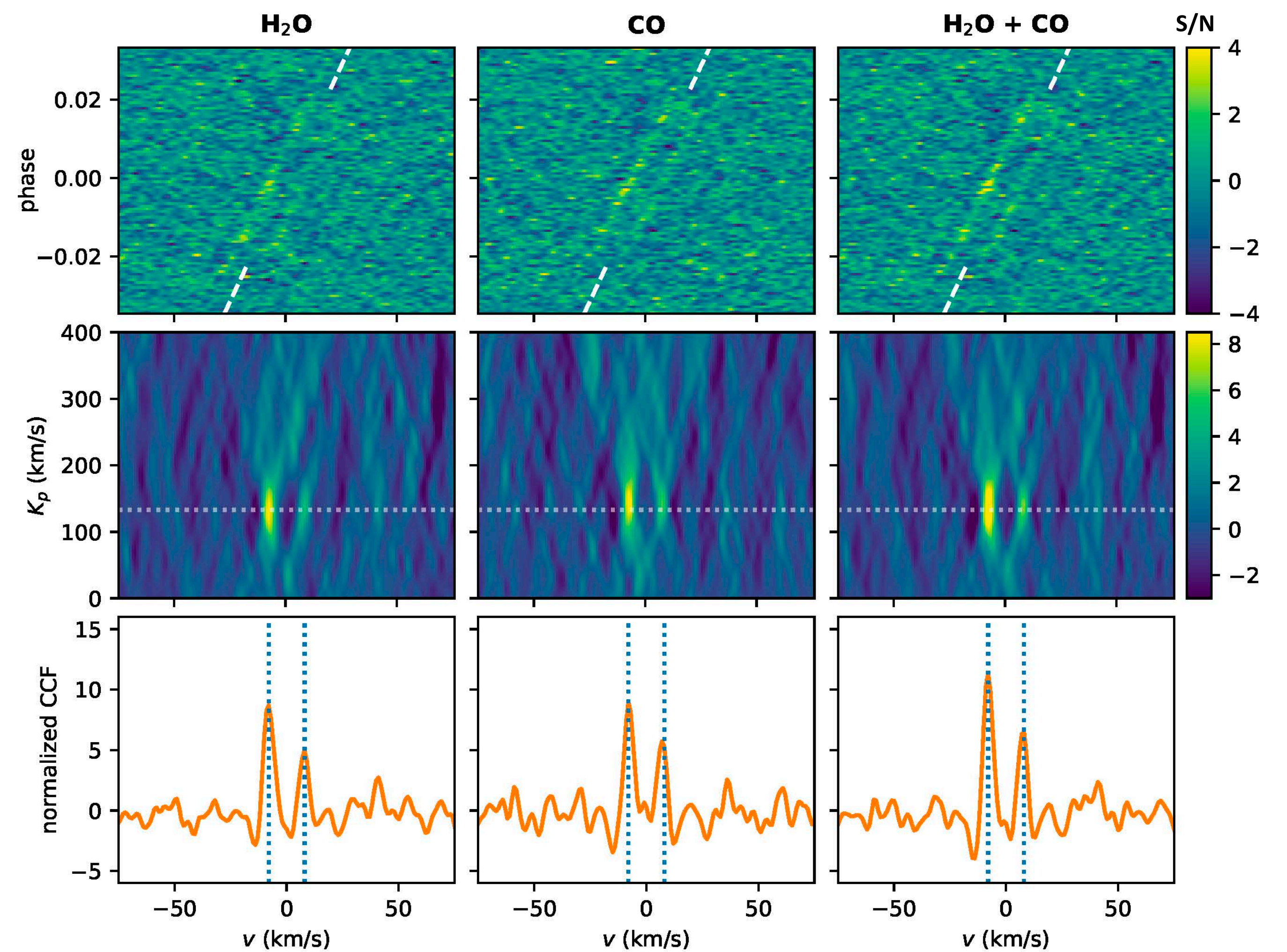
We investigated the transmission spectrum of the atmosphere of **WASP-127b** using data of one transit event observed in the K-band using **CRIRES+**. The spectral resolution achieved during the night was **R ~ 140'000**.

Using the cross-correlation method we detect **CO** and **H<sub>2</sub>O** but no **CH<sub>4</sub>** nor **CO<sub>2</sub>**. Previously CO had not been unambiguously identified in this planet<sup>11,12</sup>.

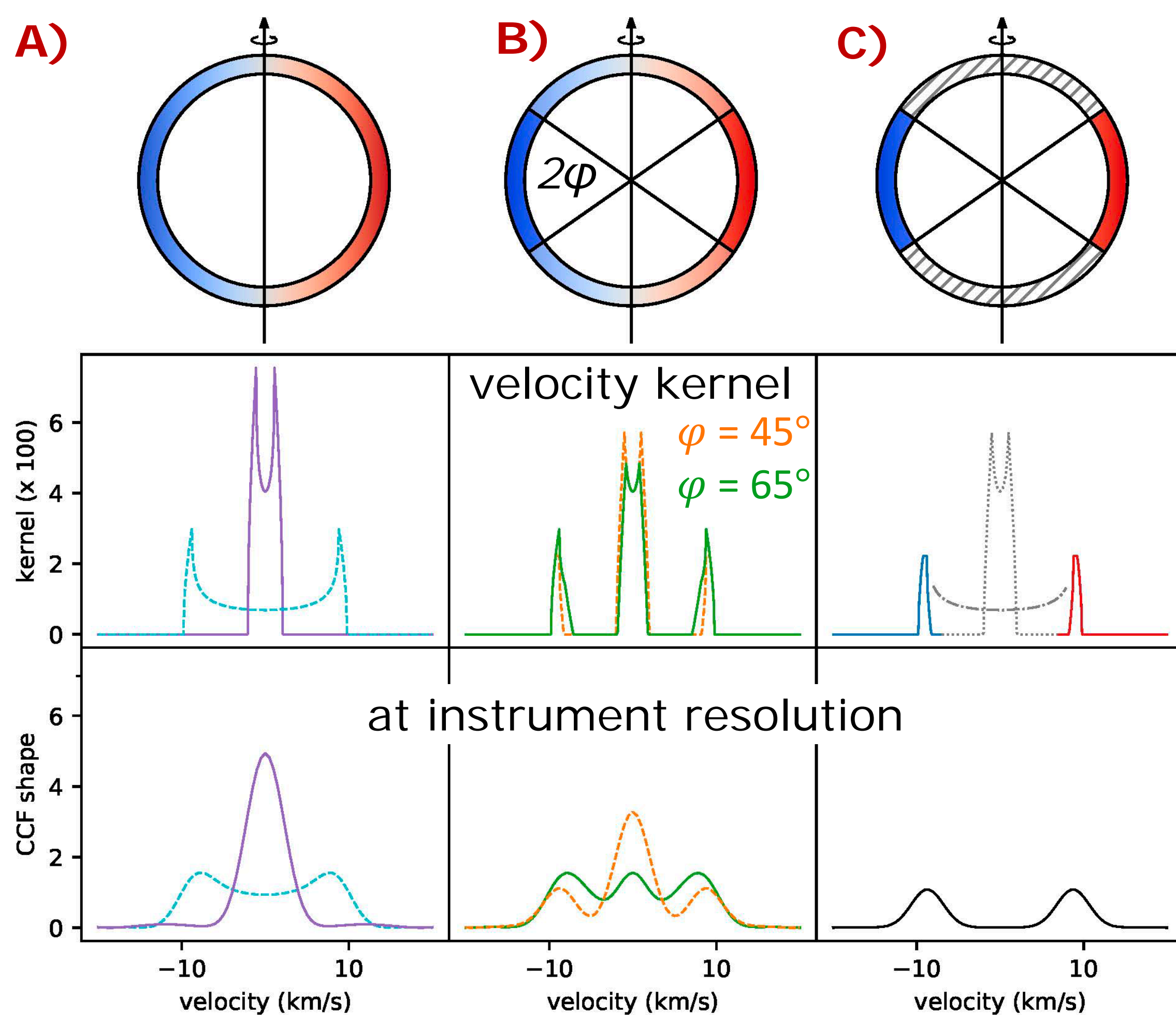
The signals show **2 resolved peaks** at  $\pm 8$  km s<sup>-1</sup>, which we can trace back to the evening and morning terminator of the planet.

We note a **lack of signal at planet rest frame**, where signals from the planetary poles would be expected.

What can explain this interesting velocity distribution of the material? Let's check the velocity kernel of a rotating atmosphere!



### Explanation of the velocity kernel



- A)** Solid body rotation → **double peaked velocity kernel**  
At low rotational velocities (e.g. tidally locked rotation), the two signals **remain unresolved at instrument resolution**. At **higher velocities** ( $\sim 10$  km s<sup>-1</sup>) the two signals can be resolved. The **plateau** between the two peaks stems from material at the poles.
- B)** If the super-rotation is constrained to an **equatorial jet**, and the polar regions rotate at tidally locked velocity, the profile has a third peak instead of a plateau.
- C)** Since we detect neither a plateau nor a third peak, we conclude that atmospheric **signals at the polar regions are likely muted**.

Let's check what a 2D retrieval says!

### Full retrieval with simplified 2D model

- Highlights:**
- **Solar C/O ratio and metallicity**, also in agreement with the values for the host star.
  - Tentatively cooler temperatures for the morning terminator than the evening terminator and even **cooler temperatures or a higher cloud deck at the polar regions**.
  - Jet extends to latitudes of  $65^\circ$  and has a **supersonic jet velocity** of  $9.3$  km s<sup>-1</sup>

