The Influence of Photometry on Deconfusion of Directly Imaged Multi-Planet Systems

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1. The Confusion Problem

- * Directly imaged multi-planet systems (with future observatories like HWO, ELTs, or Roman) are likely to introduce a "confusion" problem, making it difficult to differentiate between Earth-like and non-Earth-like planets^[1]
- * Confusion arises from not knowing which detection in a set of images belongs to which planet (see Fig. 1), due to lack of prior knowledge about orbital parameters or planetary





Fig. 1. The confusion problem from [1]. Three simulated detections of a 2-planet system over a period of 1.2 years (left). Orbits can be assigned (partitioned) in multiple different ways, resulting in substantially different orbit fits for the system (center and right).

- * The "deconfuser" was developed to address confusion by:
 - Deconfusing directly imaged planetary systems
 - Predicting confusion rates of simulated planetary systems
- * MC simulations with the deconfuser indicate higher probability of confusion for high-inclination systems^[1]. High-*i*

Fig. 2. Flow of the photometry ranking algorithm as it fits within the deconfuser.

3. Analysis of a Confused System



systems offer the greatest potential for reducing confusion with the addition of photometric considerations.

2. Deconfusion with Photometry

- * We developed a method to augment the deconfuser's ranking metrics and incorporate planetary phase information.
- * We developed a photometry model, noise model, and likelihood ranking scheme to expand the existing orbit ranking scheme
- * The flow of the expanded ranking scheme is shown in Figure 2:
 - 1. Accept orbital geometry of the planet detections as determined by the deconfuser
 - Calculate the expected flux ratio and photon count rate at the detector for each observation
 Add detector properties (e.g., read noise, dark current) to simulate noisy detection

With astrometry + photometry



- 4. Calculate likelihood of each detection given measured photometry
- 5. Calculate likelihood of each orbit
- 6. Compare orbit options

Stay tuned for Hasler et al. with full photometry analysis, coming soon!



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[1] Pogorelyuk, L., Fitzgerald, R., Vlahakis, S., Morgan, R., and Cahoy, K. (2022). ApJ, 937(2):66.
[2] Keithly, D. R. and Savransky, D. (2021). ApJL, 919(1):L11.