

Introduction

The desire to characterize and model the atmospheres of the many extrasolar planets that have been discovered over the last three decades is a major driver of current astronomy. This goal is impacted by the lack of spectroscopic data for the molecules in question since it requires significant quantities of spectroscopic data, to **accurately reproduce the spectroscopic features** of the atmospheres of hot exoplanets.

The ExoMol database provides **molecular data** for interpreting spectra and modelling atmospheres of hot exoplanets, cool stars, brown dwarfs, and other hot astronomical atmospheres and its molecules are selected because they are deemed to be important for **exoplanets and other studies of hot atmospheres**.

The core of the ExoMol database is **comprehensive high-temperature molecular line lists**; these are supplemented with **partition functions, state lifetimes, cooling functions, Landé g-factors, temperature-dependent cross sections, opacities, pressure broadening parameters, k-coefficients and dipoles**.

The ExoMol line lists are systematically published as a series in the journal Monthly Notices of the Royal Astronomical Society and summarised in release papers.

High-resolution line list database ExoMolHR [1]

The **ExoMolHR** database stores empirical high-resolution spectra extracted from the ExoMol with low uncertainty line positions.

ExoMolHR aims to precisely determine and characterize transition frequencies for high-resolution research, and this focus appears to be relevant for the updates of the ExoMol database.

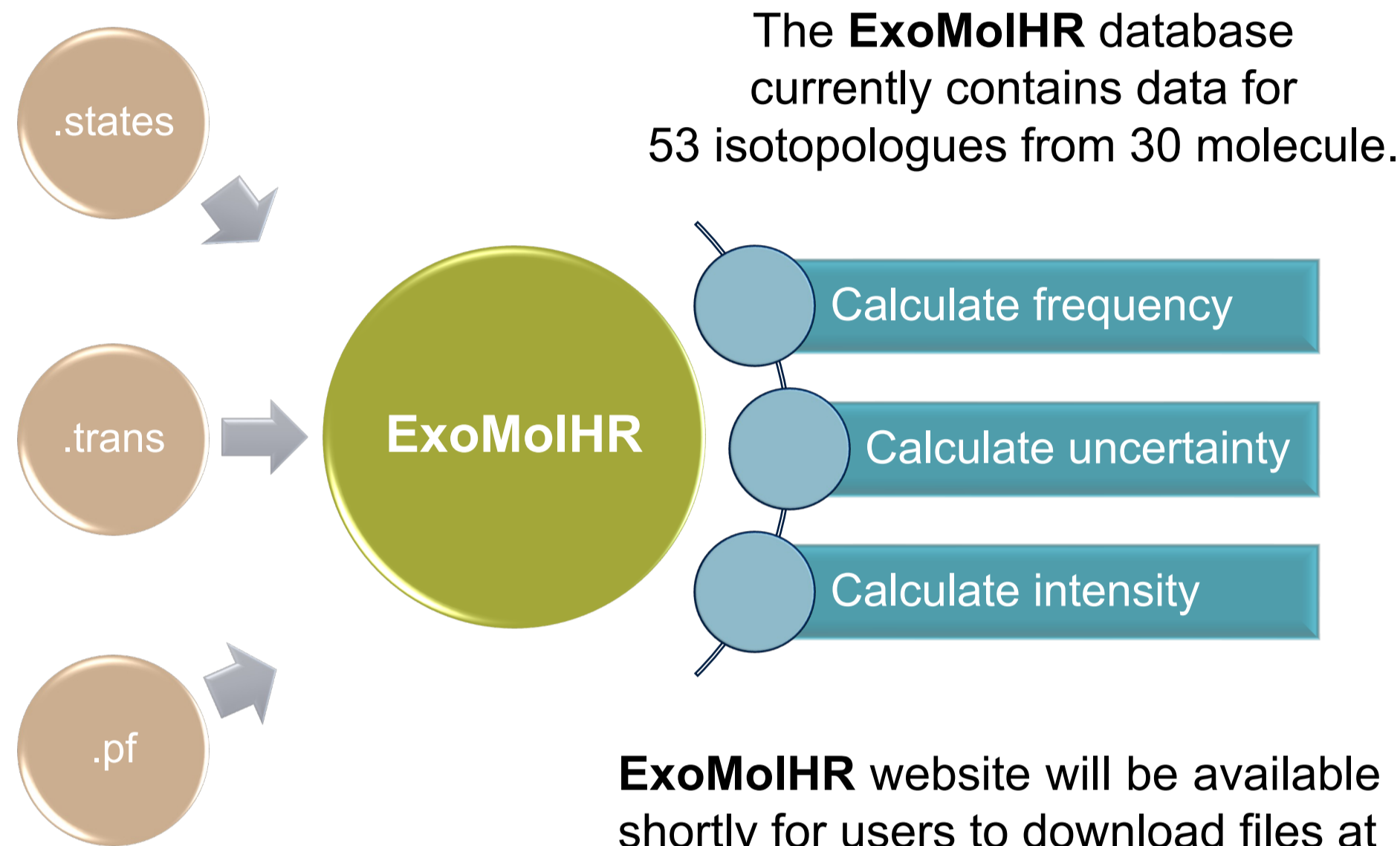


Table 1. Sample ExoMolHR results for H₂O.

Frequency	Uncertainty	A	Intensity	Molecule	Isotopologue	Dataset	E'	g'	J'	J''	QN'	QN''
0.400554	0.000002	9.35E-10	2.39E-28	H ₂ O	1H2-16O	POKAZATEL	1907.61557	27	21	4 3 2 3 0 1 0 B1	3 0 0 1 0 B2	
0.741682	0	1.96E-09	4.32E-25	H ₂ O	1H2-16O	POKAZATEL	446.510665	39	33	6 5 1 6 0 0 0 B2	2 3 0 0 0 B1	
0.895955	0.000003	7.93E-09	3.82E-28	H ₂ O	1H2-16O	POKAZATEL	2129.59892	33	27	5 4 3 2 0 1 0 B2	4 1 0 1 0 B1	
3.210931	0.000002	4.69E-07	1.74E-27	H ₂ O	1H2-16O	POKAZATEL	2126.40758	9	11	4 5 4 0 0 1 0 A1	3 3 0 1 0 A2	

Table 2. Summary of current ExoMolHR database.

Molecule	Isotopologue	Dataset	HR N Lines
AlCl	²⁷ Al ³⁵ Cl	YNAT	101
	²⁷ Al ³⁷ Cl	YNAT	121
AlH	²⁷ Al ¹ H	AloHa	692
	²⁶ Al ¹⁶ O	ATP	143197
	²⁷ Al ¹⁶ O	ATP	149577
AIO	²⁷ Al ¹⁷ O	ATP	142905
	²⁷ Al ¹⁸ O	ATP	142976
C ₂	¹² C ₂	8states	445682
C ₂ H ₂	¹² C ₂ ¹ H ₂	aCeTY	473850
CH ₄	¹² C ¹ H ₄	MM	
CN	¹² C ¹⁴ N	Trihybrid	244808
CO ₂	¹² C ¹⁶ O ₂	UCL-4000	2600218
CaH	⁴⁰ Ca ¹ H	XAB	12341
CaOH	⁴⁰ Ca ¹⁶ O ¹ H	OYT6	
H ₂ CO	¹ H ₂ ¹² C ¹⁶ O	AYTY	317729
H ₂ CS	¹ H ₂ ¹² C ³² S	MOTY	
H ₂ O	¹ H ₂ ¹⁶ O	POKAZATEL	3520554
H ₂ S	¹ H ₂ ³² S	AYT2	63719
H ₃ O ⁺	¹ H ₃ ¹⁶ O ⁺	eXeL	1785
	¹ H ₂ ² H ⁺	ST	646
	¹ H ₃ ⁺	MiZaTeP	13606
	² H ₂ ⁺ H ⁺	MiZo	683
	² H ₃ ⁺	MiZo	225
LiOH	⁶ Li ¹⁶ O ¹ H	OYT7	840
	⁷ Li ¹⁶ O ¹ H	OYT7	749
MgH	²⁴ Mg ¹ H	XAB	2462
	²⁶ Mg ¹ H	XAB	5850
	²⁸ Mg ¹ H	XAB	5339
	¹⁴ N ¹ H	kNigHt	26131
	¹⁴ N ² H	kNigHt	943
	¹⁵ N ¹ H	kNigHt	943
	¹⁵ N ² H	kNigHt	943
NH ₃	¹⁴ N ¹ H ₃	CoYuTe	412149
NO	¹⁴ N ¹⁶ O	XABC	106711
NaOH	²³ Na ¹⁶ O ¹ H	OYT5	
SO	³² S ¹⁶ O	SOLIS	2501
SO ₂	³² S ¹⁶ O ₂	ExoAmes	1504495
	²⁸ Si ¹⁴ N	SiNfull	670
	²⁸ Si ¹⁵ N	SiNfull	464
	²⁹ Si ¹⁴ N	SiNfull	464
	³⁰ Si ¹⁴ N	SiNfull	464
SiO	²⁸ Si ¹⁶ O	SiOUVenIR	8729
TiO	⁴⁸ Ti ¹⁶ O	Toto	499775
VO	⁵¹ V ¹⁶ O	HyVO	
	⁸⁹ V ¹⁶ O	BRYTS	25
	⁸⁹ V ¹⁷ O	BRYTS	25
	⁸⁹ V ¹⁸ O	BRYTS	25
	⁹⁰ Zr ¹⁶ O	ZorrO	145317
	⁹¹ Zr ¹⁶ O	ZorrO	5164
	⁹² Zr ¹⁶ O	ZorrO	5164
	⁹³ Zr ¹⁶ O	ZorrO	5164
	⁹⁴ Zr ¹⁶ O	ZorrO	5164
	⁹⁶ Zr ¹⁶ O	ZorrO	5164

Developments

Continuous updating of the ExoMol database is particularly important.

We present the recent developments of the ExoMol database (<https://www.exomol.com/>).

Adopting the IAEA standards

for the atomic and molecular information

Table 3. Extract from the states file of the ²⁷AlH line list.

i	E (cm ⁻¹)	g	J	unc (cm ⁻¹)	r (s)	+/	off	State	r	IAI	IBI	Scale (cm ⁻¹)	Source Label
1	0.000000	12	0	0.000000	inf	-	e	XI(Sigma+)	0	0	0	0.000000	Ma
2	1625.961581	12	0	0.000000	4.9208E-03	+	e	XI(Sigma+)	1	0	0	1625.969321	Ma
3	3194.213550	12	0	0.000000	2.6227E-03	+	e	XI(Sigma+)	2	0	0	3194.213585	Ma
4	4708.888866	12	0	0.000000	1.8633E-03	+	e	XI(Sigma+)	3	0	0	4708.877022	Ma
5	6170.189966	12	0	0.000000	1.4919E-03	+	e	XI(Sigma+)	4	0	0	6170.193041	Ma
6	7579.564189	12	0	0.000000	1.2740E-03	+	e	XI(Sigma+)	5	0	0	7579.564189	Ca
7	8938.646805	12	0	0.000000	1.1350E-03	+	e	XI(Sigma+)	6	0	0	8938.646805	Ca
8	10386.646827	12	0	0.000000	1.0410E-03	+	e	XI(Sigma+)	7	0	0	10386.646827	Ca
9	11506.245734	12	0	0.000000	9.7520E-04	+	e	XI(Sigma+)	8	0	0	11506.245734	Ca
10	12717.633797	12	0	0.000000	9.2834E-04	+	e	XI(Sigma+)	9	0	0	12717.633797	Ca

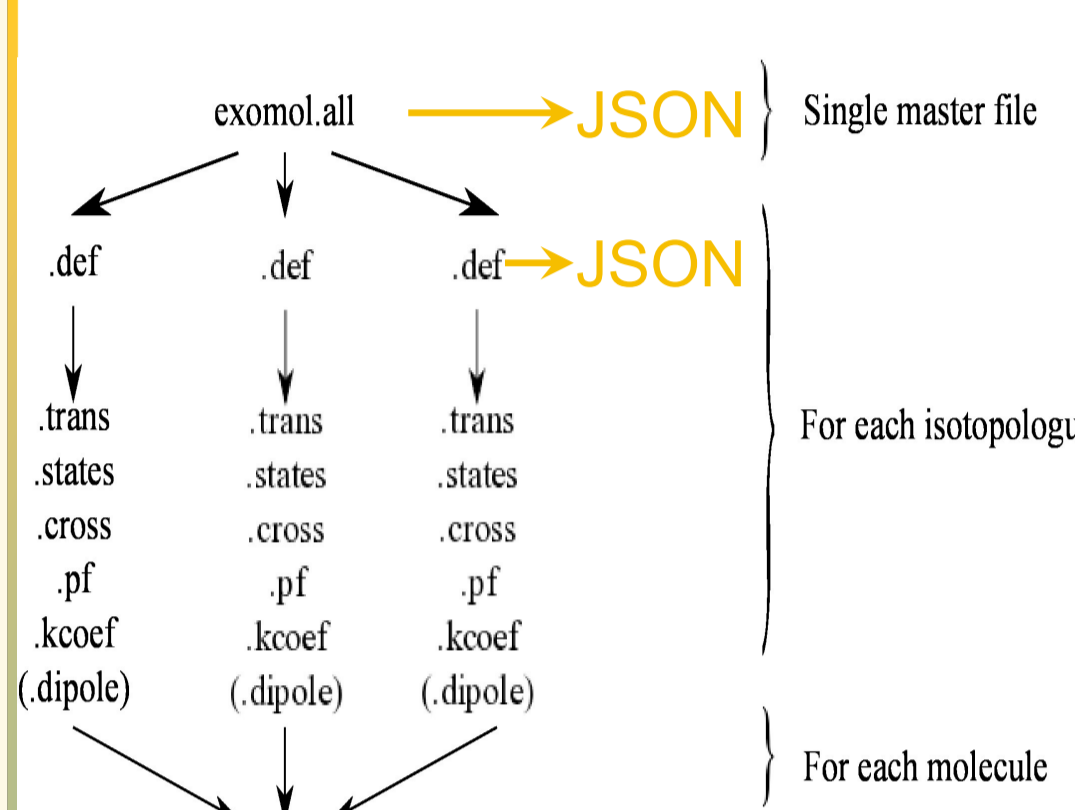
We have updated our quantum number specifications to conform with the **PyValem** python package for parsing, validating, and manipulating quantum states labels of atoms, ions, and small molecules.

Ancillary database LiDB [2]

Molecular vibronic state radiative lifetimes for plasma processes

<https://www.exomol.com/liadb/>

JSON format for line list APIs



Photodissociation cross sections and continuum absorption

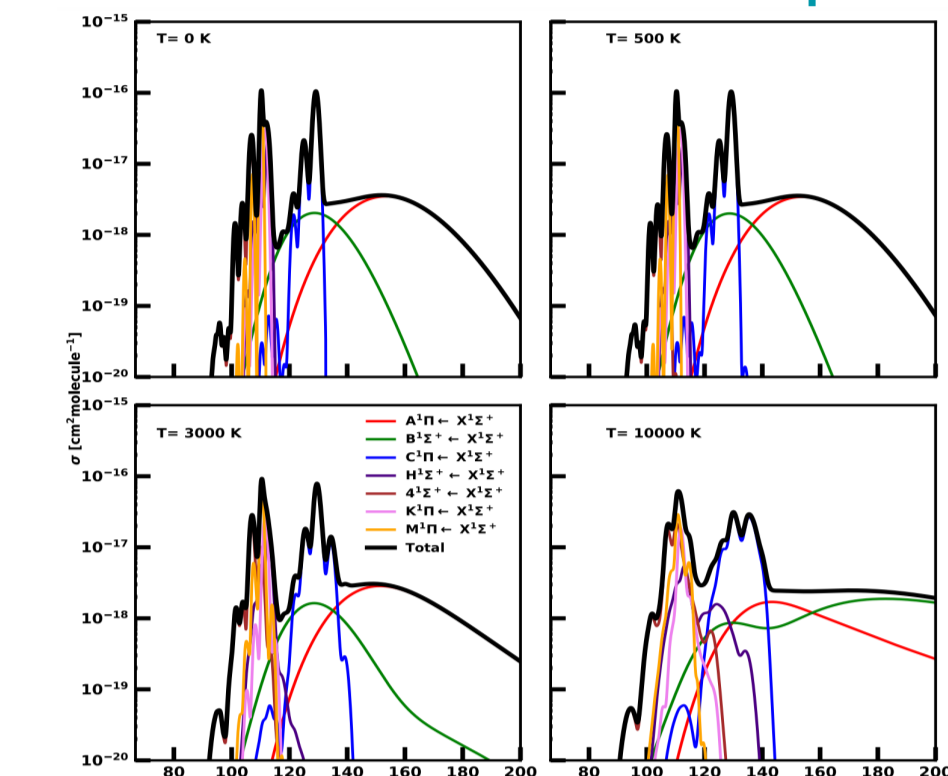


Figure 1. Total and the partial photodissociation cross-sections of H³⁵Cl at 0, 500, 3000, and 10000 K [3]. We investigate the temperature dependence of the photodissociation cross-sections and rates for molecules of importance to exoplanetary.

High accuracy line lists with inclusion of uncertainties

I. BeH, MgH, CaH	XI. HNO ₃	XXI. NO	XXXI. C ₂	XLI. NaOH, KaOH
II. SiO	XII. CS	XXII. SiH ₄	XXXII. MgO	XLII. NO (UV)
III. HCN/HNC	XIII. CaO	XXIII. PO, PS	XXXIII. TiO	XLIII. NaO
IV. CH ₄	XIV. SO ₂	XXIV. SiH	XXXIV. PH	XLIV. SiO (UV)
V. NaCl, KCl	XV. HOOH	XXV. SiS	XXXV. NH ₃	XLV. MgH (UV), CaH (UV)
VI. PN	XVI. H ₂ S	XXVI. SN, SH	XXXVI. SH (UV)	XLVI. SiN
VII. PH ₃	XVII. SO ₃	XXVII. C ₂ H ₄	XXXVII. HCCH	XLVII. CaOH
VIII. H ₂ CO	XVIII. VO	XXVIII. AlH	XXXVIII. SiO ₂	XLVIII. H ₂ CS
IX. AlO	XIX. H ₂ ¹⁸ O, H ₂ ¹⁷ O	XXIX. CH ₃ Cl	XXXIX. CO ₂	XLIX. AlCl
X. NaH	XX. H ₃ ⁺	XXX. H ₂ ¹⁶ O	XL. H ₃ O ⁺	L. H ₃ ⁺ , H ₂ D ⁺ , D ₂ H ⁺ , D ₃ ⁺

High accuracy line lists
High accuracy in progress

Old line lists
In progress
O₂, NiH, CO, AlF, CH, HCO⁺, OH⁺

LI. LiOH
LII. CH⁺
LIII. YO
LIV. AlD
LV. VO (hyperfine)
LVI. SO
LVII. CH₄
LVIII. OCS
LIX. N₂O
LX. ¹⁵NH₃
LXI. OH
LXII. C₃
LXIII. HDO
LXIV. PN

A Python program PyExoCross [4] is designed to post process both ExoMol and HITRAN line lists

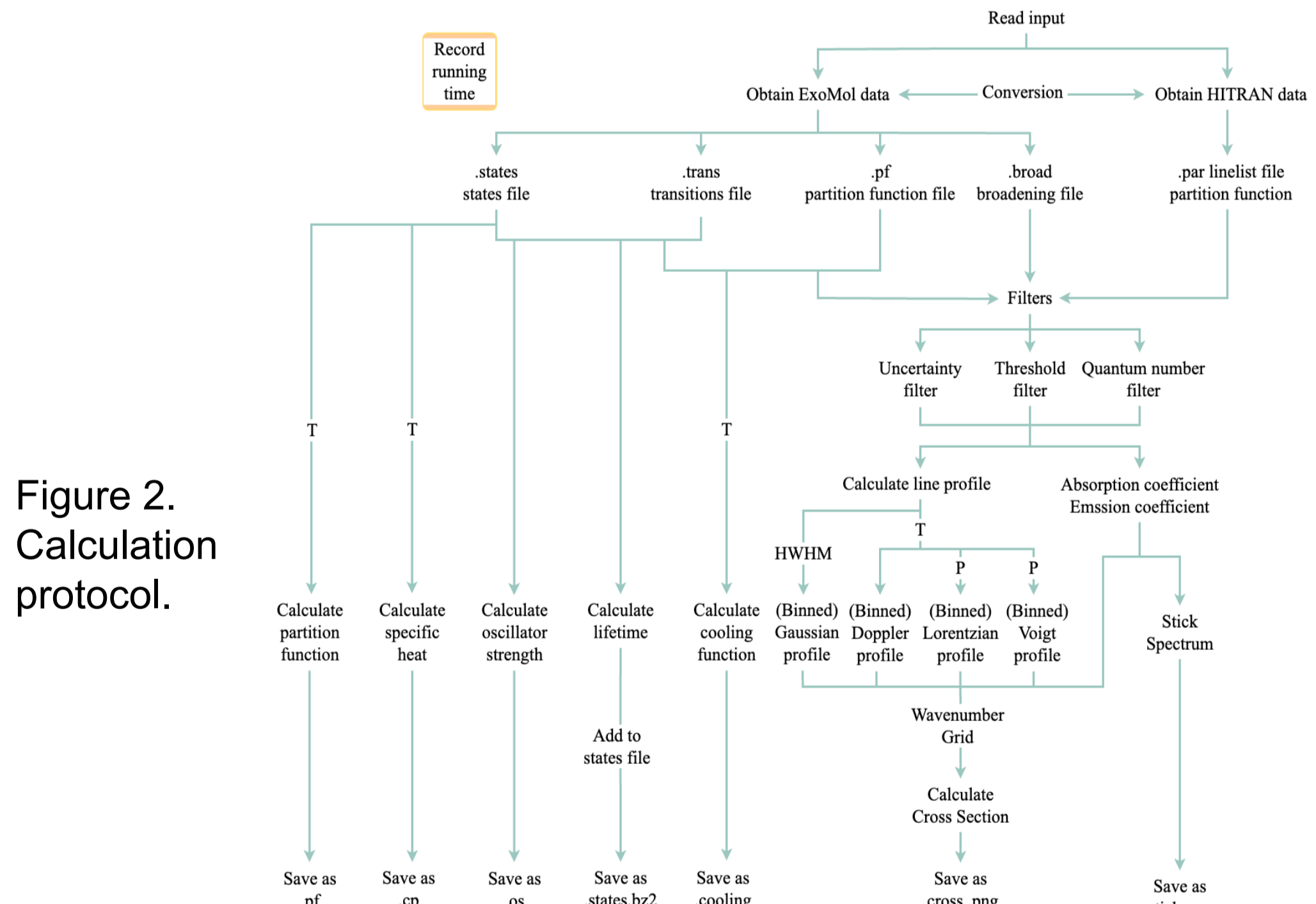


Figure 2. Calculation protocol.

Figure 3(a). Partition functions.

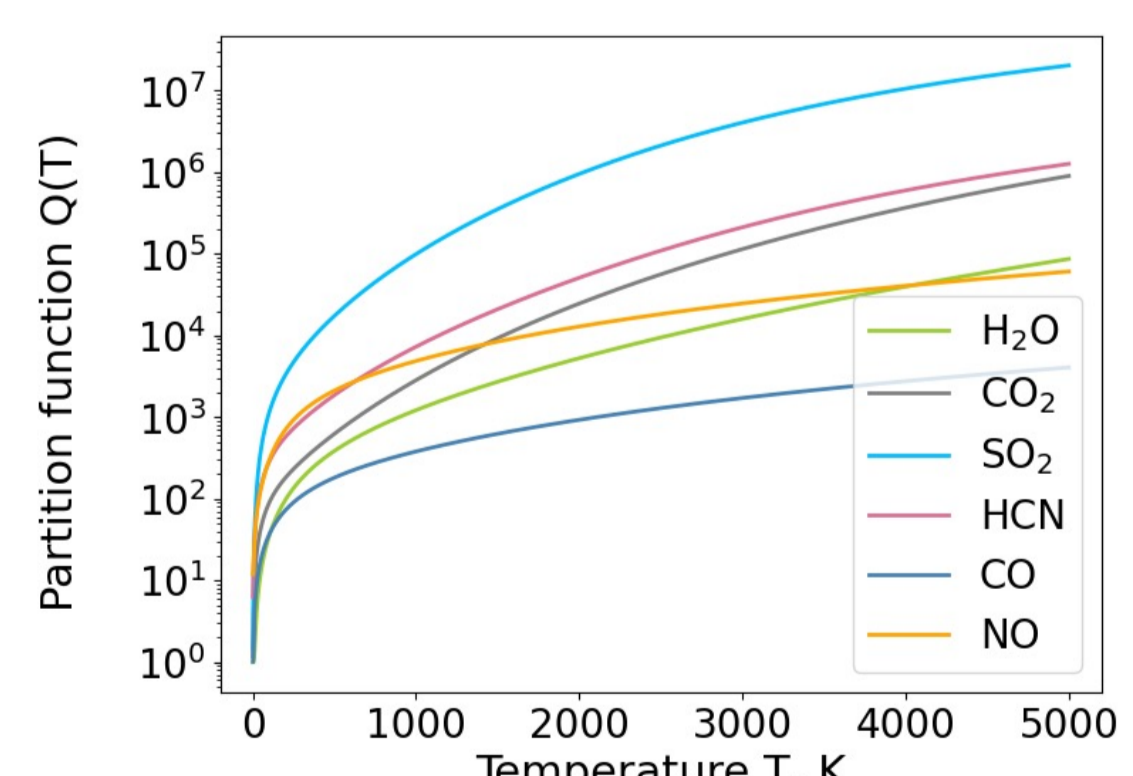


Figure 3(c). Cooling functions.

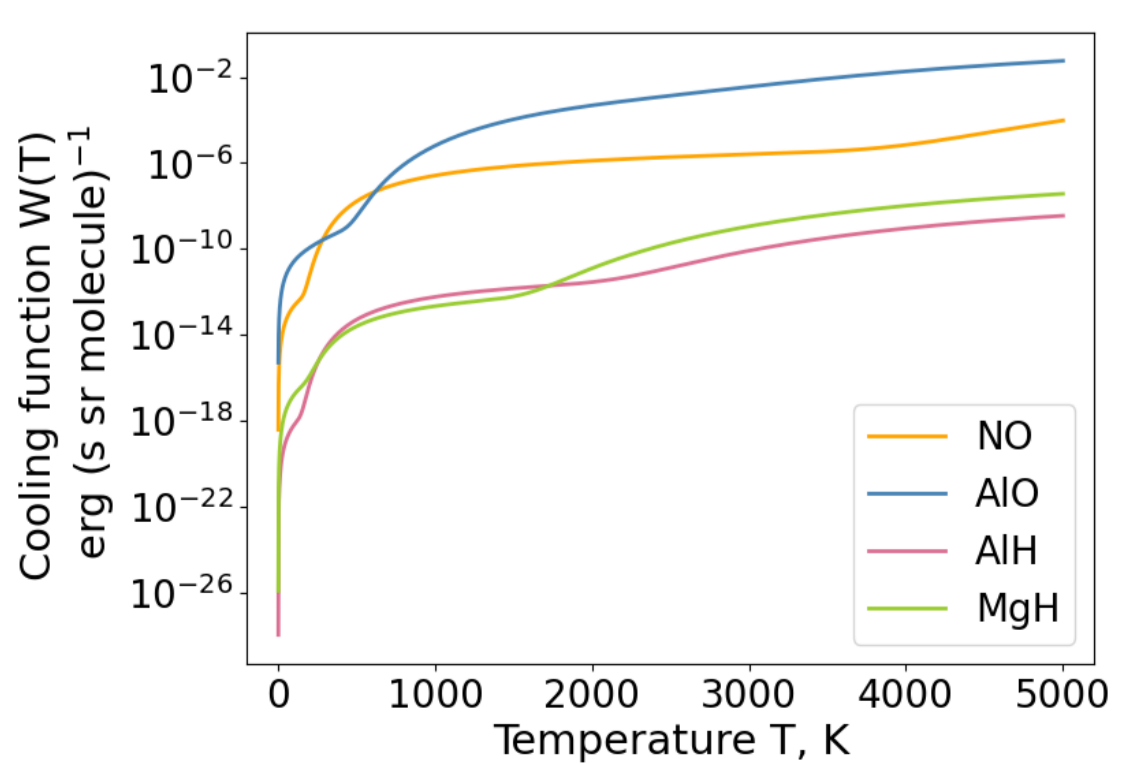


Figure 3(b). Specific heats.

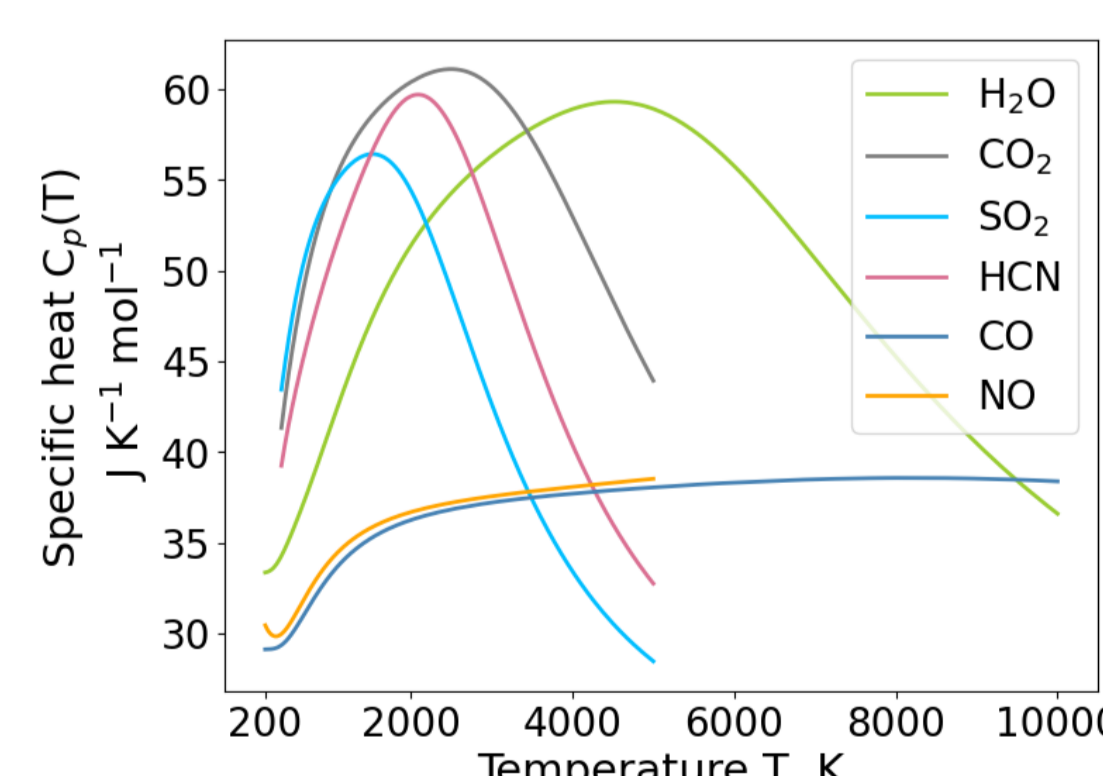


Figure 3(d). Lifetimes.

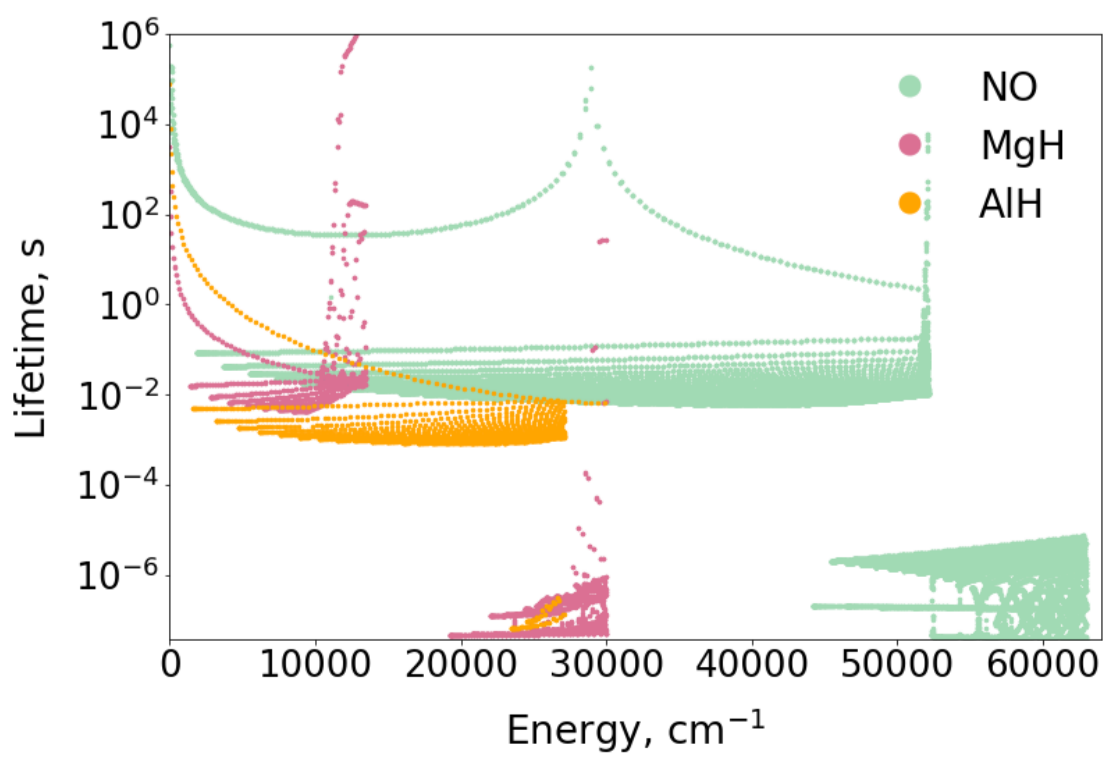


Figure 3(e). Oscillator strength.

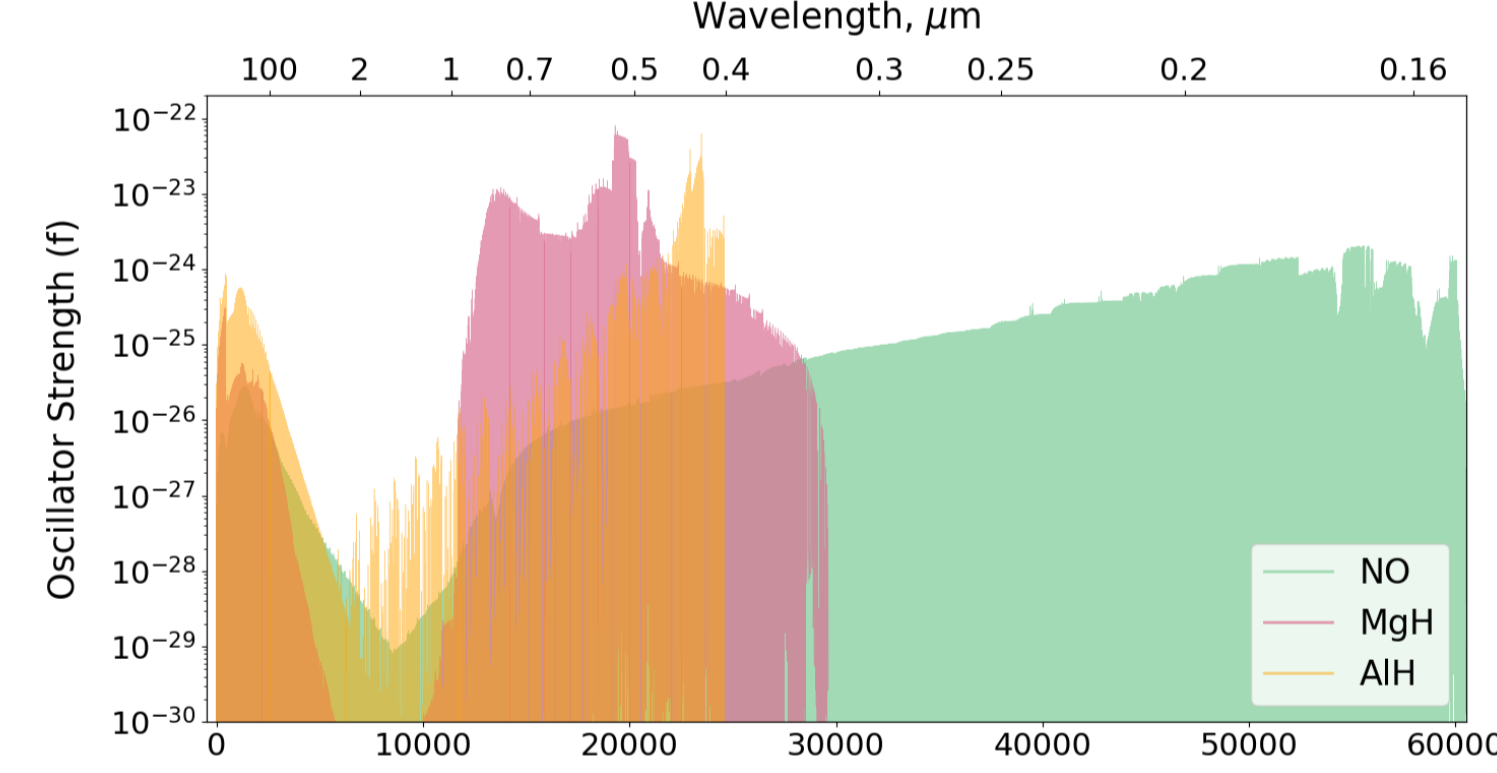


Figure 3(f). Stick spectra at different temperatures.

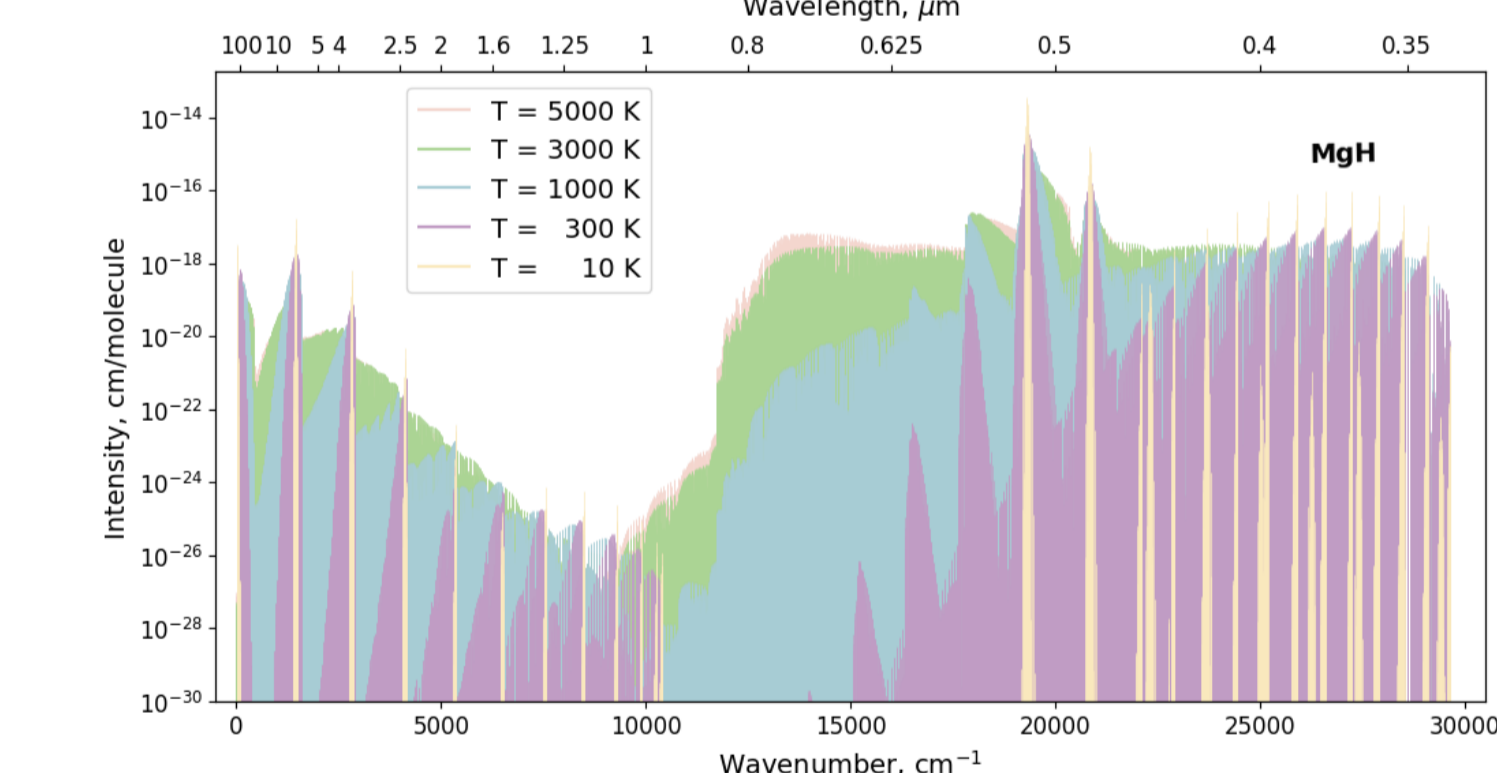


Figure 3(g). Stick spectra for different electronic states.

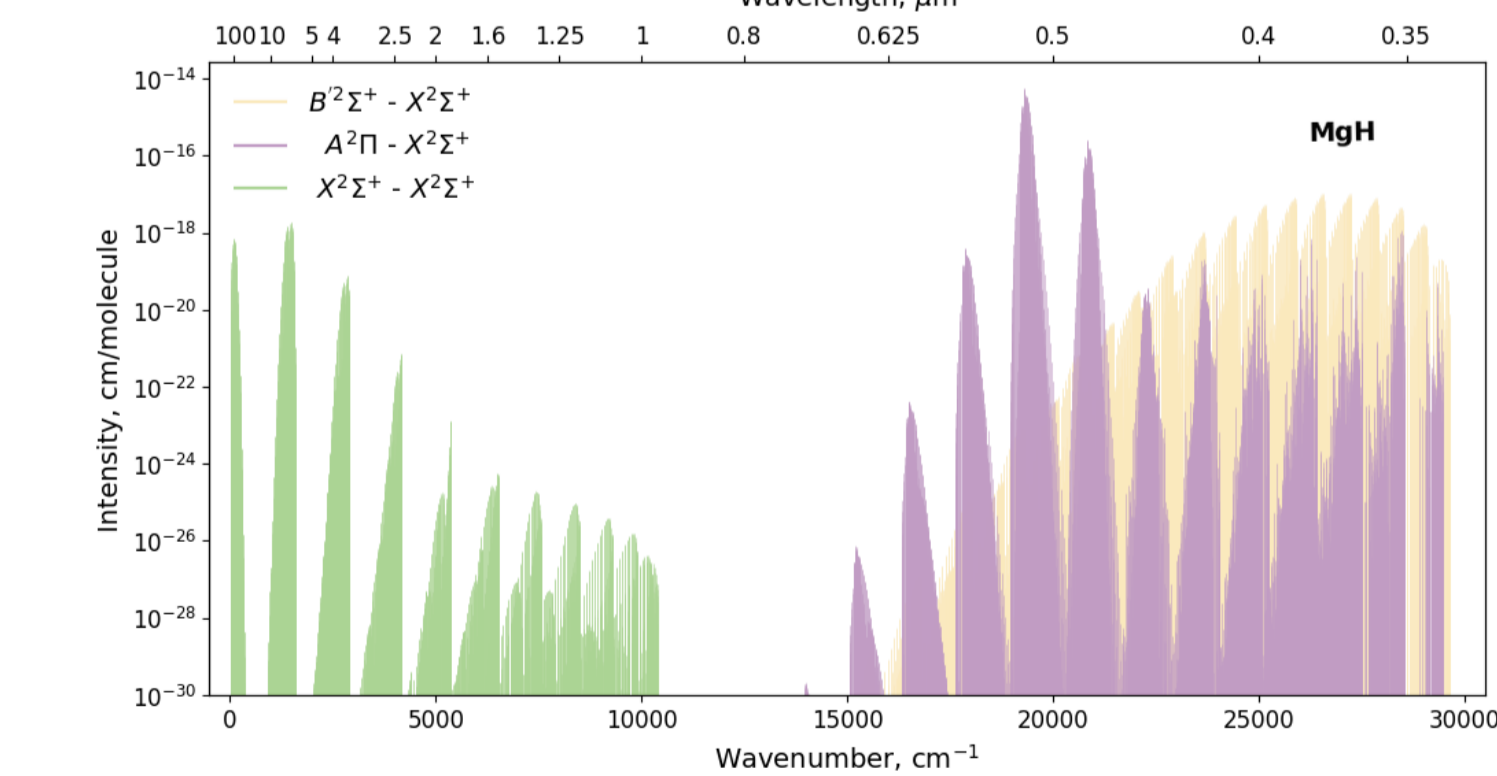
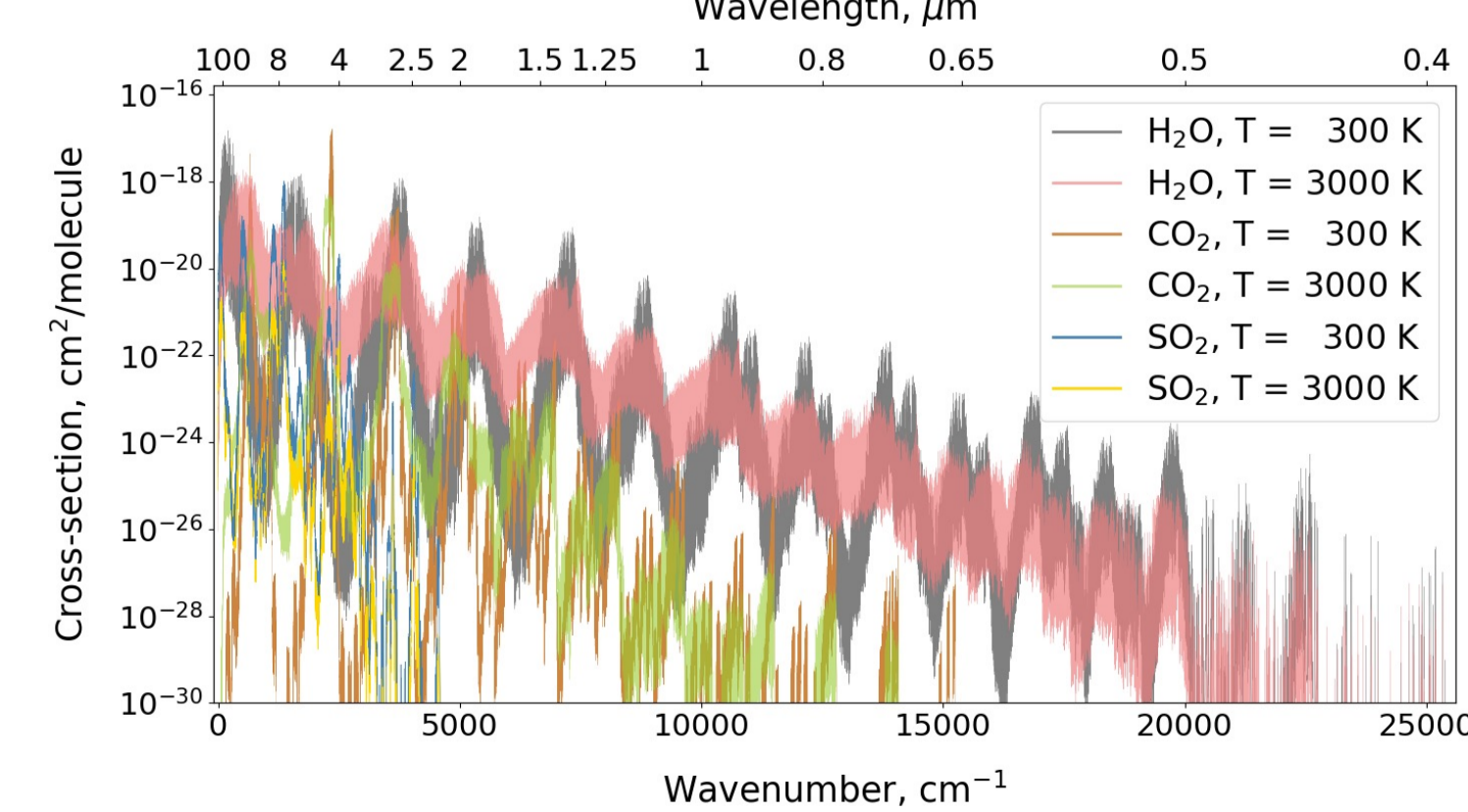


Figure 3(h). Cross sections at different temperatures.



Finding PostDoc Position

Interests: Database, Molecular Spectroscopy, Exoplanet, Cross Sections
Programming Languages: Python, Linux, Fortran, Django, HTML
Skills: Scientific Computing, Data Mining, Machine Learning, Web Crawler, Data Analysis, Data Visualization

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<https://github.com/Beryl-Jingxin>

[1] Jingxin Zhang et al., "ExoMolHR: A Relational Database of Empirical High-Resolution Molecular Spectra", in preparation (2024).

[2] A. Owens et al., *Plasma Sources Sci. Tech.*, **32**, 085015 (2023).

[3] M. Pezzella et al., *Mon. Not. R. astr. Soc.*, **514**, 4413-4425 (2022).

[4] Jingxin Zhang et al., "PyExoCross: a Python program for generating spectra and cross sections from molecular line lists", *RAS Tech. Instr.*, (2024).

